

Computation of Earth Rotation Parameters Consistent with the IERS Earth Rotation Representation

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IERS Terrestrial (T) to Celestial (C) Coordinate Transformation:

$$\mathbf{R} = \mathbf{Q}_0(X, Y) \mathbf{R}_3(s(X, Y)) \mathbf{R}_3(-\theta(\tau)) \mathbf{R}_3(-s'(x_p, y_p)) \mathbf{R}_2(x_p) \mathbf{R}_1(y_p)$$

Geocentric bases: $\bar{\mathbf{e}}^c = [\bar{e}_1^c \ \bar{e}_2^c \ \bar{e}_3^c]$ $\bar{\mathbf{e}}^t = [\bar{e}_1^t \ \bar{e}_2^t \ \bar{e}_3^t]$

Position vector: $\bar{\mathbf{x}} = \bar{\mathbf{e}}^c \mathbf{x}_c = \bar{\mathbf{e}}^t \mathbf{x}_t$

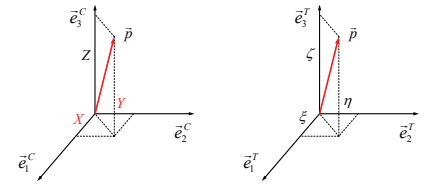
Base & coordinate transformation: $\bar{\mathbf{e}}^c = \bar{\mathbf{e}}^t \mathbf{R}^T$ $\mathbf{x}_c = \mathbf{R} \mathbf{x}_t$

Involves CIP (Celestial Intermediate Pole):

$$\bar{\mathbf{p}} = \bar{\mathbf{e}}^c \mathbf{p}_c = \bar{\mathbf{e}}^t \mathbf{p}_t, \quad |\bar{\mathbf{p}}| = 1, \quad \mathbf{p}_c = \begin{bmatrix} X \\ Y \\ Z \end{bmatrix}, \quad \mathbf{p}_t = \begin{bmatrix} \xi \\ \eta \\ \zeta \end{bmatrix} \approx \begin{bmatrix} x_p \\ -y_p \\ \zeta \end{bmatrix}$$

with 5 Transformation Parameters: $X, Y, x_p, y_p, \tau = \text{UT1} - \text{UTC}$ 2 more than 3 required for defining the orthogonal matrix $\mathbf{R} \rightarrow$ **Inconsistency!**

The CIP is determined from theoretical solutions to the earth rotation problem and removal of frequencies beyond the observational resolution



The (mathematically) Compatible Rotation Vector:

$$\bar{\mathbf{e}}^t = \bar{\mathbf{e}}^c \mathbf{R} \Rightarrow \frac{d\bar{\mathbf{e}}^t}{dt} = \bar{\mathbf{e}}^c \frac{d\mathbf{R}}{dt} = \bar{\mathbf{e}}^t \mathbf{R}^T \frac{d\mathbf{R}}{dt} = [\omega_t \times] \bar{\mathbf{e}}^t, \quad \bar{\omega} = \bar{\mathbf{e}}^t \omega_t = \bar{\mathbf{e}}^c \omega_c, \quad \omega_c = \mathbf{R} \omega_t$$

$$[\omega_c \times] = \frac{d\mathbf{R}}{dt} \mathbf{R}^T, \quad [\omega_t \times] = \mathbf{R}^T \frac{d\mathbf{R}}{dt}, \quad \omega \equiv |\bar{\omega}| = \sqrt{\omega_c^T \omega_c} = \sqrt{\omega_t^T \omega_t}$$

The (mathematically) Compatible Celestial Pole (CCP):

$$\bar{\mathbf{n}} = \frac{\bar{\omega}}{\omega} = \bar{\mathbf{e}}^c \mathbf{n}_c = \bar{\mathbf{e}}^t \mathbf{n}_t, \quad \mathbf{n}_c = \frac{1}{\omega} \omega_c \equiv \begin{bmatrix} \bar{X} \\ \bar{Y} \\ \bar{Z} \end{bmatrix}, \quad \mathbf{n}_t = \frac{1}{\omega} \omega_t \equiv \begin{bmatrix} \bar{\xi} \\ \bar{\eta} \\ \bar{\zeta} \end{bmatrix} \approx \begin{bmatrix} \bar{x}_p \\ -\bar{y}_p \\ \bar{\zeta} \end{bmatrix}, \quad \mathbf{R} = \mathbf{Q}_0(\bar{X}, \bar{Y}) \mathbf{R}_3(s(\bar{X}, \bar{Y})) \mathbf{R}_3(-\theta(\bar{\tau})) \mathbf{R}_3(-s'(\bar{x}_p, \bar{y}_p)) \mathbf{R}_2(\bar{x}_p) \mathbf{R}_1(\bar{y}_p)$$

The CCP is determined from observations only, it is mathematically compatible with the rotation matrix \mathbf{R} and it is independent from any earth rotation theory

Computation of the CCP = Computation of "Observed" Precession-Nutation, Polar Motion and Length of the Day: (Dermanis, 2005, Journées des Systèmes de Référence Spatio-Temporels, Warsaw)

$$\begin{bmatrix} \bar{X} \\ \bar{Y} \end{bmatrix} = \frac{\partial}{\partial \omega} \begin{bmatrix} X \\ Y \end{bmatrix} + \frac{1}{\omega} \begin{bmatrix} 1 - aX^2 & -aXY \\ -aXY & 1 - aY^2 \end{bmatrix} \begin{bmatrix} b_0^1 \\ b_0^2 \end{bmatrix} - \begin{bmatrix} \cos \psi & -\sin \psi \\ \sin \psi & \cos \psi \end{bmatrix} \begin{bmatrix} b_w^1 \\ b_w^2 \end{bmatrix}, \quad \begin{bmatrix} \bar{\xi} \\ \bar{\eta} \end{bmatrix} = \frac{\partial}{\partial \omega} \begin{bmatrix} \xi \\ \eta \end{bmatrix} + \frac{1}{\omega} \begin{bmatrix} 1 - a'\xi^2 & -a'\xi\eta \\ -a'\xi\eta & 1 - a'\eta^2 \end{bmatrix} \begin{bmatrix} b_w^1 \\ b_w^2 \end{bmatrix} + \begin{bmatrix} \cos \psi & \sin \psi \\ -\sin \psi & \cos \psi \end{bmatrix} \begin{bmatrix} b_0^1 \\ b_0^2 \end{bmatrix}, \quad \psi = \theta - s + s', \quad a = \frac{1}{1 + \sqrt{1 - X^2 - Y^2}}, \quad a' = \frac{1}{1 + \sqrt{1 - \xi^2 - \eta^2}}$$

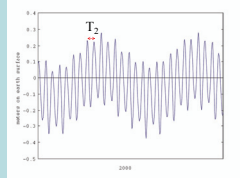
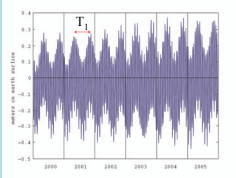
$$\begin{bmatrix} b_0^1 \\ b_0^2 \end{bmatrix} = \frac{a^2}{2a-1} \begin{bmatrix} X(Y\bar{X} - X\bar{Y}) - \frac{aY}{1-a}(X\bar{X} + Y\bar{Y}) \\ Y(Y\bar{X} - X\bar{Y}) + \frac{aX}{1-a}(X\bar{X} + Y\bar{Y}) \end{bmatrix}, \quad \begin{bmatrix} b_w^1 \\ b_w^2 \end{bmatrix} = \frac{a^2}{2a-1} \begin{bmatrix} \xi(\eta\bar{\xi} - \xi\bar{\eta}) - \frac{a'\eta(\xi\bar{\xi} + \eta\bar{\eta})}{1-a'} \\ \eta(\eta\bar{\xi} - \xi\bar{\eta}) + \frac{a'\xi(\xi\bar{\xi} + \eta\bar{\eta})}{1-a'} \end{bmatrix}, \quad \mathbf{R} = \mathbf{Q}_0(\bar{X}, \bar{Y}) \mathbf{R}_3(s(\bar{X}, \bar{Y})) \mathbf{R}_3(-\theta(\bar{\tau})) \mathbf{R}_3(-s'(\bar{x}_p, \bar{y}_p)) \mathbf{R}_2(\bar{x}_p) \mathbf{R}_1(\bar{y}_p)$$

solve for $\theta(\bar{\tau})$
 $\theta(\tau) = A + B \cdot \text{UT1} = A + B(\text{UTC} + \tau) \Rightarrow \bar{\tau} = \text{UT1} - \text{UTC} = \frac{\theta(\bar{\tau}) - A}{B} - \text{UTC}$

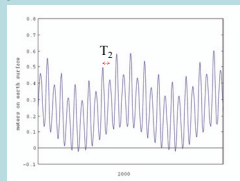
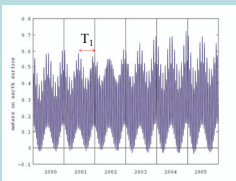
Comparison of Observed Earth Rotation (CCP = Compatible Celestial Pole) with Theoretical Earth Rotation (CIP = Celestial Intermediate Pole)

Precession-Nutation

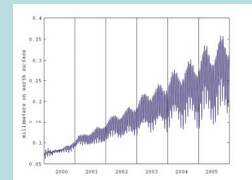
$$X_{\text{CCP}} - X_{\text{CIP}}$$



$$Y_{\text{CCP}} - Y_{\text{CIP}}$$



$$S_{\text{CCP}} - S_{\text{CIP}}$$



Two periodic terms with periods

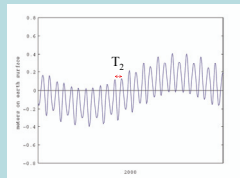
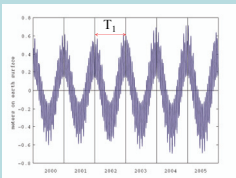
$T_1 = 190.5$ days
(\approx semiannual)

$T_2 = 13.6$ days
(\approx half lunar month)

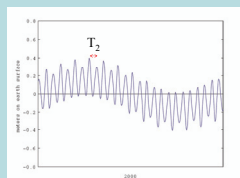
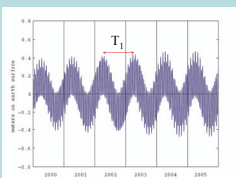
negligible differences

Polar Motion

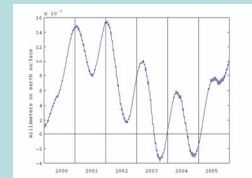
$$(x_p)_{\text{CCP}} - (x_p)_{\text{CIP}}$$



$$(y_p)_{\text{CCP}} - (y_p)_{\text{CIP}}$$



$$s'_{\text{CCP}} - s'_{\text{CIP}}$$



Two periodic terms with periods

$T_1 = 357.5$ days
(\approx annual)

$T_2 = 14.2$ days
(\approx half lunar month)

negligible differences

Length of the Day

$$\tau = \text{UT1} - \text{UTC}$$

$$\tau_{\text{CCP}} = \tau_{\text{CIP}}!$$

Identical!